

Lava Tubes





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The Moon has been extensively visited, since the very beginning of the space era, by remote sensing techniques from orbiting satellites, in-situ analysis from landed bases, etc. Furtehrmore, it is the only celestial body visited by human crews :

5 Fly-by	12 Impact (-2)	5 Probes
2 Test flights (-2)	11 Landers (-3)	19 Orbiters
3 Sample Return	1 Crewed Orbiter	7 Crewed Landers (-1)

2 Rovers

1959-2007: 66 (+1) missions attempted (8 aborted/failed)





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Such a large number of missions have produced a large amount of data about our natural satellite, providing an overall knowledge of its nature and phenomena.

In the mid of '70s, the moon exploration program was discontinued (similarly to what happened to the Mars exploration). A stop which lasted until the '90s, when new missions have been implemented.

The new missions have been focused on specific investigations, also in view of a possible manned return to the moon.

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Clementine (NASA, 1994): Cartography, Composition.

An experiment obtained data (at the poles) compatible with water

Main Payload: Laser Altimeter, IR, VIS&UV cameras

Lunar Prospector (NASA, 1998): Composition, Magnetic properties. Further evidence of water ice.

Main Payload: Magnetometer, γ-spectrometer





SMART-1 (ESA, 2003): Technology mission. Topography, Morphology, Geology, water in craters

Main Payload: IR Spectrometer, X-Ray Spectrometer, Microcamera

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SELENE (Kaguya) (JAXA, 2007): Origin, Geology, Environment

Main Payload: HR Camera, Laser Altimeter, Magnetometer, Radar sounder





Global coverage, including poles

Main Payload: Stereo Camera, Laser Altimeter, Imaging-/gamma-/X-ray spectrometers, microwave radiometer, High energy particles detector





Moon Exploration



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Courtesy M. A. Perino



In the frame of the manned exploration of the Moon, some precursor missions have to be implemented, which shall be focused on the acquisition of further usefull information to be used to decide upon the location and design the human habitat.

Investigation (jointly performed by scientific and industrial communities) have been activated in Italy on various aspects of the moon exploration and, in particular, on the information needs and of the relevant gathering techniques.

An example of information of possible important use is the location and characteristics of the lava tubes

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Lava Tubes on Earth....

Lava tubes are formed by a flowing low-viscosity lava stream. It develops a hard crust, which becomes the roof covering the still flowing lava. As a final result an empty buried channel is left.

These structures are common on Earth.



Malpais National Monument, New Mexico – The entrance of a lava tube

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PAVONIS MONS

Trundolo

... and on Mars ...





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... and on the Moon



Lunar Orbiter Image V-182-M

(USGS, 1987)

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Lava Tubes for Lunar Bases

Several authors support the idea of utilizing the lava tubes on the Moon to build the first lunar base.

In fact with their roofs, which are tens of meter thick, they represent a natural shield against cosmic radiation and solar flares, as well as against ejecta from meteorite impacts.



Under lunar conditions of low gravity and absence of atmosphere, lava tubes are espected to be much larger than on Earth (tens/hundreds of meters in cross section and tens of kilometers in length)

Mapping of lava tubes, (location and size) can be an important mission objective

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Sounding Radar



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Sounding Radar



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Sounding Radars on Mars

Mars Express (ESA, 2003): MARSIS

MRO (NASA, 2005): ShaRad



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Echo power level in color scale. Surface returns are stroger (deep red), potential subsurface echoes are weaker and at different range

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ShaRad vs. MARSIS



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ShaRad Radar Gram



Example of SHARAD radargram relevant to an area with several subsurface layers

Further to the mapping/characterization of lava tubes, the processing of the Radar Sounders echoes can provide information on interesting geological properties (e.g. thickness of regolith layer)

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Mapping Lava Tubes with GPR

SUBSURFACE GEOPHYSICAL IMAGING OF LAVA TUBES, LAVA BEDS NATIONAL MONUMENT, CA

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Figure 7 – GPR data collection at Hercules with the 400 MHz antenna and survey wheel



Figure 11 – GPR profile over Hercules Leg Cave (known caves are outlined in blue while an additional anomaly is outlined red)



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Modeling Lava Tubes - 1



- Having a known geometry, the Lava Tube echo to a sounding radar can be properly modeled and therefore detected
 Moreover analyzing the electromagnetic characteristic and correlating the echoes coming from the Moon surface, the upper and lower part of the tube, is possible to understand the physical composition of:
 - ✓ the tube cavity (e.g. if empty or not)
 - ✓ the tube perimeter
 - ✓ the rock between the Lava Tube and Moon surface



Being a sounder nadir looking, the typical observation geometries of a lava tube can be parallel, orthogonal or diagonal with respect to the ground track. Depending on that, the radargram will present different features



and green part of the cylinder); the visible portion of the tube which will depend on the ratio between the radar wavelength and the tube diameter (D)

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 $\frac{2\mathbf{D}}{2\mathbf{D}} \cdot \sqrt{\mathbf{\varepsilon}_r \mathbf{\mu}_r}$

 $\frac{2S}{2} \cdot \sqrt{\varepsilon_r \mu_r}$



Modeling Lava Tubes - 3







Buried Crater found by Marsis Instrument

In this type of geometry, the radargram will show the rings (known phenomenon observed by MARSIS and SHARAD sounders when flying over craters). Intensity of these rings will be function of the inclination (maximum at 45°, close to zero when parallel or orthogonal)





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Radar Sounder: preliminary requirements

In the frame of Italian Vision for Moon Exploration Program funded by Italian Space Agency, an evaluation of a Radar Sounder on a lunar orbiter with the objective to map the lava tubes has been performed.

An assessment of the preliminary radar parameters and possible instrument architecture has been carried out.



In order to optimize the main instrument parameters (carrier frequency, signal bandwidth, power, etc.), a deeper modeling of lunar subsurface dielectric characteristics will be necessary

Main Parameters		
Carrier	100 MHz	
Bandwidth	25 MHz	
PRF	700 Hz	
Vertical Resolution	6 m	
Range Resolution	2.2 Km	
Azimuth resolution	50 m	
Max Penetration Depth	0.3 Km	
Pulse Width	75 μsec	
Peak TX Power	20 W	
DC Power	55 W	
Mass	13 Kg	
Size	0.40(L)x0.45(W) x0.2(H) mxmxm	
Antenna Length	1.5 m	
Nbit	12	
Data rate	46 8 Mbits/sec	



 A complete map of lunar lava tubes is an important information in view of the exploration and colonization of the Moon

- As lava tubes are subsurface structures, they are difficult to be identified on surface images and, in any case, a catalogue obtained only by surface images would inevitably be incomplete
- The potential of low frequency radar technology to identify lava tubes has been demonstrated on Earth applications
- Radar sounder technology for planetary exploration can be considered at a maturity stage and is presently successfully in operation in Mars orbit
- A small lunar orbiter, embarking a radar sounder for lava tubes mapping, can be based on existing (or easily derivable) technologies, with obvious advantages in economical and programmatic terms







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